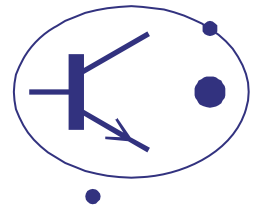


# BRITISH PHYSICS OLYMPIAD

<http://www.BPhO.org.uk>



**Please note that this mark scheme is a final draft version. It contains several small errors that were corrected during the marking process.**

# SOLUTIONS

# BPhO 2010 PAPER 2

Q1

MARK

(a)

(i) Energy lost in 8.4 s =  $5 \times 140 \times 3$  J

1

Rate of loss of energy from lead =  $5 \times 140 \times 3 / 8.4$  J s<sup>-1</sup>  
 = 250 W

1

(ii) If specific latent heat of fusion is  $L$  then

$$3L = 250 \times 300$$

$$L = \underline{2.5 \times 10^4 \text{ J kg}^{-1}}$$

1

(iii) If specific heat capacity of liquid lead is  $s$  then

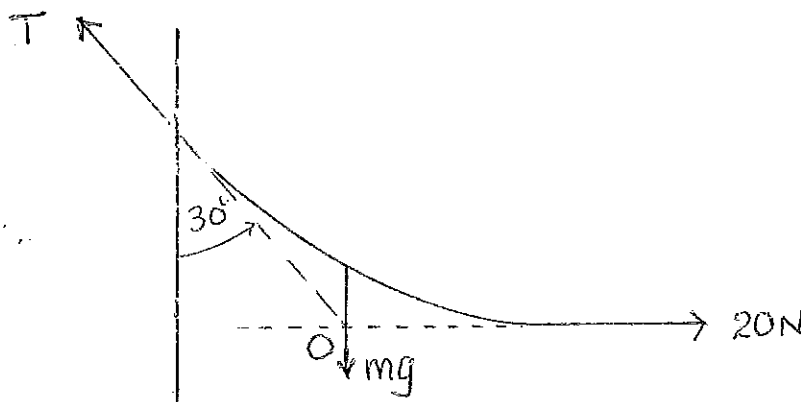
$$3(5)s = 250 \times 10$$

$$s = \underline{167 \text{ J kg}^{-1} \text{ K}^{-1}}$$

1

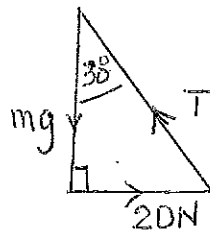
[4]

(b)



Vertical weight of rope must pass through O

Triangle of forces



1

$$\tan 30^\circ = \frac{20}{mg}$$

$$mg = 20 \cot 30^\circ$$

1

$$= 20\sqrt{3}$$

$$= 34.64 \text{ N}$$

$$m = \underline{3.53 \text{ kg}}$$

1

[3]

(c) Surface area of a bit =  $\frac{\pi (5.5^2 - 2.2^2)}{650 \times 10^6 \times 8} = \frac{79.8}{5.2 \times 10^9} \text{ cm}^2$  2  
 $= 1.5 \times 10^{-8} \text{ cm}^2 = 1.5 \times 10^{-12} \text{ m}^2$  1  


---

 3

(d) Area of lens =  $\pi (10^{-2})^2 = \pi 10^{-4} \text{ m}^2$   
 Energy pers (1% of 100W) =  $\frac{1}{100} 100 = 1 \text{ W}$   
 Time =  $1.5 \times 10^{-2} \text{ s}$   
 Energy reaching film =  $\frac{\pi (10^{-4})}{4\pi (10^2)^2} (1) 1.5 \times 10^{-2}$  2  
 $= 3.8 \times 10^{-11} \text{ J}$  1

For  $n$  photons of frequency  $\nu$   
 $n h \nu = 3.8 \times 10^{-11}$   
 As  $\nu = \frac{c}{\lambda}$   
 $n (6.6 \times 10^{-34}) \frac{3.00 \times 10^8}{6 \times 10^{-7}} = 3.8 \times 10^{-11}$  1  
 $n = \frac{3.8 \times 10^{-11}}{3.3 \times 10^{-19}}$   
 $n = 1.2 \times 10^8$  (Accept  $1.1 \times 10^8$ ) 1  

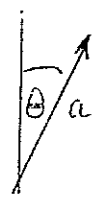

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 [5]

(e) At temperature  $T \text{ K}$   
 $\left(\frac{250}{230}\right)^2 20 \left[ 1 + (T-273) 5.0 \times 10^{-3} \right] = 250$  2  
 $\frac{(230)^2}{250} \frac{1}{20} = 1 + (T-273) 5.0 \times 10^{-3}$  } 1  
 $T = 273 + \frac{(230)^2}{25} \frac{1}{5.0 \times 10^{-3}}$   
 $T = 2.2 \times 10^3 \text{ K}$  1  


---

 [4]

(f) For amplitude  $a$  incident at angle  $\theta$   
 Transmitted amplitude =  $a \cos \theta$   
 This has a transmitted intensity =  $a^2 \cos^2 \theta$   1  
 Average over all  $\theta$ , =  $\langle a^2 \cos^2 \theta \rangle = \frac{1}{2} a^2$  1  
 Adding over all amps. for all frequencies in white light  
 intensity =  $\frac{1}{2} \sum_i a_i^2 = \frac{1}{2} I$  as  $I = \sum_i a_i^2$  1  


---

 [3]

(g) (i) To escape from planet with initial velocity  $v$  requires the KE. to be at least equal to initial gravitational P.E.

$$\frac{1}{2}mv^2 = \frac{GMm}{R} \quad 2$$

$$v = \sqrt{\frac{2GM}{R}} \quad 1$$

(ii) For planets to have no atmosphere  $\langle |v| \rangle$  must be much greater than  $\sqrt{\frac{GM}{R}}$  so that they escape from gravitational field

ie.  $\langle |v| \rangle \gg \sqrt{\frac{GM}{R}}$

A factor greater than 5 is sufficient

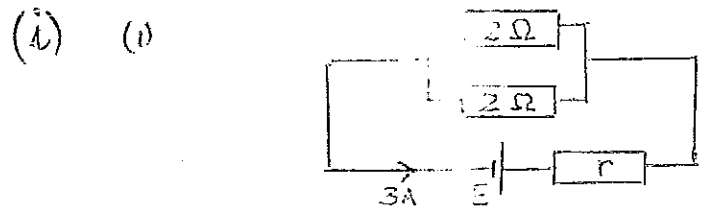
[4]



Mirror rotates through 1' of arc  
 Time taken =  $1/(35 \times 360 \times 60)$   
 Distance travelled = 400m

Velocity of light  $c = 400 \times 35 \times 360 \times 60 \text{ ms}^{-1}$   
 $c = 3.0 \times 10^8 \text{ ms}^{-1}$

[3]

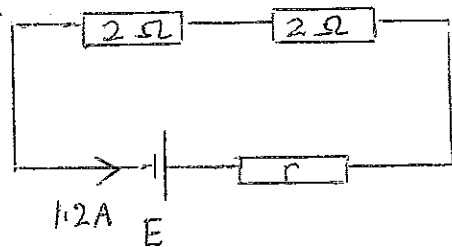


$$E = 3 \left( r + \left( \frac{1}{2} + \frac{1}{2} \right)^{-1} \right)$$

$$= 3(r+1)$$

$$r = \frac{E}{3} - 1$$

1



$$E = 1.2(r + 4) \quad \frac{1}{2}$$

Sub<sup>g</sup> from earlier result

$$E = 1.2 \left( \frac{E}{3} - 1 + 4 \right) \quad \frac{1}{2}$$

$$= 0.4E + 3.6$$

$$\underline{E = 6V} \quad \frac{1}{2}$$

$$(i) \quad r = \frac{E}{3} - 1 = \frac{6}{3} - 1 = 1\Omega \quad \frac{1}{2}$$

$$(ii) \quad P_p = \left(\frac{3}{2}\right)^2 2 = 4.5W \quad \frac{1}{2}$$

$$P_s = (1.2)^2 2 = 2.88W \quad \frac{1}{2}$$

[4]

$$(i) \text{ Initial thrust of rocket} = 16 \times 2.5 \times 10^3 \text{ N} = 4 \times 10^4 \text{ N} \quad \frac{1}{2}$$

This is less than weight of rocket =  $9.81 \times 4.1 \times 10^3 = 4.03 \times 10^4 \text{ N}$   
 Consequently rocket does not lift off immediately.

$$(ii) \text{ Thrust of rocket is constant} = 4 \times 10^4 \text{ N}$$

Weight of rocket reduces with time

$$\text{After time } t \text{ mass of rocket} = 4.1 \times 10^3 - 16t \quad \frac{1}{2}$$

$$\text{Weight of rocket} = 9.81(4.1 \times 10^3 - 16t) \quad \frac{1}{2}$$

$$\text{Lift off occurs when} \quad 4 \times 10^4 = (4.1 \times 10^3 - 16t) 9.81 \quad 1$$

$$\therefore 16t = 4.1 \times 10^3 - \frac{4 \times 10^4}{9.81} = 22.5$$

$$\underline{t = 1.4S} \quad \frac{1}{2}$$

[4]

(k) (i)  $\frac{1}{2}$  mark for any two suggestion, for example:

No interactions between molecules

Ideal gas equations

Zero volume of all molecules

Random motion of molecules.

Mean KE  $\propto$  absolute temperature ( $\langle KE \rangle = \frac{3}{2}kT$ )

1

(ii)  $\frac{1}{2}$  mark for any two suggestions for example:

Real gas molecules interact

Gas equations deviate from ideal gas laws

Motion no longer purely random

Molecules have finite volume

$\langle KE \rangle \neq \frac{3}{2}kT$

1

(iii)  $\frac{1}{2}$  mark for any two suggestions such as:

Ideal gas satisfies  $PV = RT$

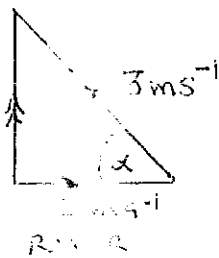
Real gas deviates from this relation

(Special cases: Boyle's & Charles laws)

Thermodynamic properties deviate from ideal values

1  
[3]

(l) (i) RESULTANT MOTION

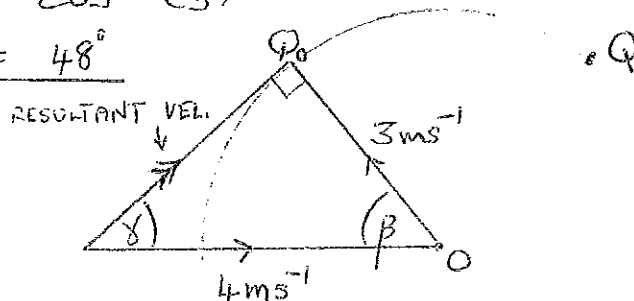


To travel perpendicular to bank he has to row at angle  $\alpha$  where

$$\alpha = \cos^{-1}\left(\frac{2}{3}\right)$$

$$\alpha = 48^\circ$$

(ii) CIRCLE CENTRE O



1  
1

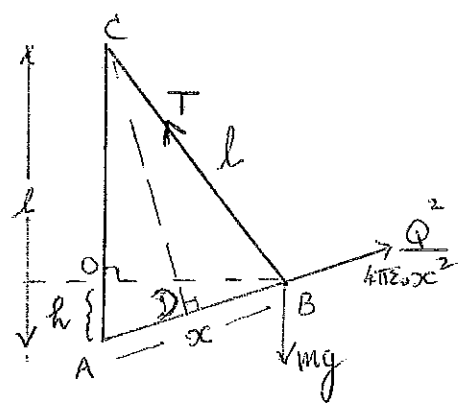
(l)

No solution possible perpendicular to bank is same as velocity of river  
 Possible directions of resultant motion along PQ, where  
 Q is on a circle, centre O, radius  $QO = 3 \text{ms}^{-1}$  in vector diagram  
 The shortest path occurs for largest value of angle  $\gamma$  which  
 will occur when PQ tangential to circle at P.  
 (Lower half circle excluded as it is on land)

$\therefore \sin \gamma = \frac{3}{4}$   
 $\gamma = \sin^{-1} \frac{3}{4} = 48.6^\circ$

1  
 2  
 1  
 1  
[7]

(m)



Isosceles triangle of forces ABC sides  $l, l$  and  $x$ .  
 Forces  $T, mg$  and  $\frac{Q^2}{4\pi\epsilon_0 x^2}$ , where  $T = mg$  as ABC isosceles.

(i)  $T = mg$

(ii)  $W = mgh + \frac{Q^2}{4\pi\epsilon_0 x}$  ①

From  $\Delta$  of forces  $\frac{x}{l} = \frac{Q^2}{4\pi\epsilon_0 x^2 mg}$  ②

In similar triangles OAB and CAD, where D mid point AB.

$\frac{h}{x} = \frac{x}{l}$

$x^2 = 2hl$  ③

Sub<sup>g</sup> ② into ①  $\frac{x}{l} = \frac{Q^2}{(4\pi\epsilon_0) 2hlmg}$  ④

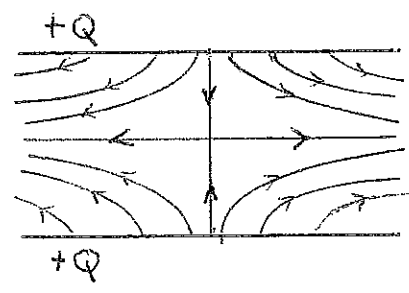
Sub<sup>g</sup> for  $x$  from ③ into ①

$W = mgh + 2mgh = 3mgh$

1  
[8]

(a)

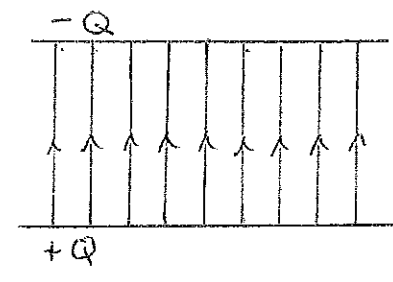
(i)



Drawing  
Correct arrows

1  
1

(ii)



Drawing  
Correct arrows

1/2  
1/2

(iii) Magnitude of charges same in both cases, (i) and (ii) and distances between all pairs of charges same in both cases. Consequently only difference, in (ii) from (i), is sign of charges. Thus if force of repulsion in (i) is  $F_+$  Force of attraction in (ii) is  $F_- = -F_+$  (using inverse square law of force between charges)

Thus 
$$\frac{F_+}{F_-} = -1$$

1

(iv) 
$$C = \frac{\epsilon_0 A}{x}$$
  
Thus energy  $E = \frac{Q^2 x}{2\epsilon_0 A}$  (A)

1  
1

(v) Force of attraction 
$$F = \frac{dE}{dx} = \frac{Q^2}{2\epsilon_0 A}$$

Alternatively comparing with gravitational energy of mass  $m$ ,  $mgh$ , which produces a force  $mg$  i.e. " $E = Fx$ "  
Thus from (A)

$$F = \frac{Q^2}{2\epsilon_0 A}$$

} 1

[7]



(o)

Let  $f$  be natural frequency and natural wavelength  $\lambda$ ,  
Wavelength  $\lambda_s$  produce by source moving towards observer, velocity  $v_s$ ,  
is given by

$$\lambda_s = \lambda - \frac{v_s}{f}$$

Then, for speed of sound  $c$ ,

$$f_s = \frac{c}{\lambda_s} = \frac{c}{\lambda - \frac{v_s}{f}}$$
$$= \frac{c}{\frac{c}{f} - \frac{v_s}{f}}$$

$$\underline{f_s = f \left( \frac{c}{c - v_s} \right)}$$

Max speed  $v_s = 2\pi(1.2)^3 = 22.6 \text{ ms}^{-1}$

Min speed  $-v_s = -22.6 \text{ ms}^{-1}$

Thus  $f$  is in the range given by

$$f_s = 256 \left( \frac{340}{340 \mp 22.6} \right)$$

Thus the variation is 240 to 274 Hz

[7]

(p)

(i) Velocity along the field  $v_{||} = v \cos \theta$

This remains constant as no force acts in this direction

Velocity perpendicular to field  $v_{\perp} = v \sin \theta$

The field  $B$  produces a constant force on the electron perpendicular to  $v_{\perp}$ , which causes it to rotate in a circle, with constant angular velocity, in this plane.

(ii) This force is  $Bev \sin \theta$  (note constant value)

For motion in a circle radius  $r$

$$\frac{m_e (v \sin \theta)^2}{r} = Bev \sin \theta$$

$$r = \frac{m_e v \sin \theta}{Be}$$

(p) Hence one period  $T$  given by

$$T = \frac{2\pi r}{v \sin \theta}$$

$$= \frac{2\pi m_e v \sin \theta}{Be v \sin \theta}$$

$$\underline{T = \frac{2\pi m_e}{Be}}$$

(iii) Distance travelled in time  $T$ ,  $L$ , is given by

$$L = T(v \cos \theta)$$

$$\underline{L = \frac{2\pi m_e v \cos \theta}{Be}}$$

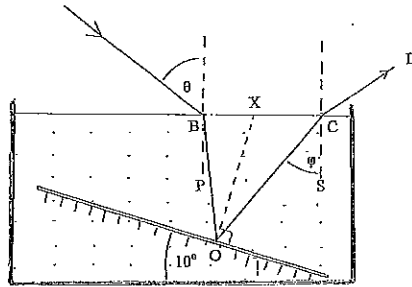
[8]

2 (a) Refractive index  $\mu = \frac{\text{REAL HEIGHT}}{\text{APPARENT HEIGHT}} = \frac{8}{5} = 1.6$

2

(b) If light emerges at C we require for glass of refractive index  $n$   
 $n \leq \frac{\sin 90^\circ}{\sin \phi} = \frac{1}{\sin \phi}$  (not totally internally reflected)

$\sin \phi \leq \frac{1}{n} = \frac{3}{4}$   
 $\phi \leq 48.59^\circ$



From diagram,

$\angle OCB = 90^\circ - \phi$

$\angle OXC = 90^\circ + 10^\circ = 100^\circ$   
 as mirror tilted by  $10^\circ$

In  $\Delta COX$

$\angle COX = \phi - 90^\circ$

$\angle BOX = \phi - 90^\circ$

angle of incid. = angle of reflec.

In  $\Delta BOX$

$\angle OBX = 110 - \phi$

angles of  $\Delta$  add to  $180^\circ$

$\therefore \angle PBO = 90^\circ - (110 - \phi) = \phi - 20$

Refraction at B:

$n = \frac{4}{3} = \frac{\sin \theta}{\sin(\phi - 20)}$

$\sin(\phi - 20) = \frac{3}{4} \sin \theta$

If  $\phi < 48.58^\circ$

$\sin(\phi - 20) \leq \sin(28.58^\circ)$

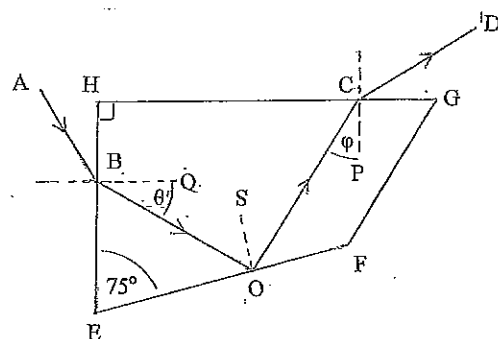
$\therefore \sin \theta \leq \frac{4}{3} \sin(28.59^\circ)$

$\theta \leq 39.6$

Maximum  $\theta$  is  $39.6^\circ$

[6]

(c)



$$\begin{aligned}
 \text{(i)} \quad \angle EBO &= 90 - \theta \\
 \angle EOB &= 180 - (90 - \theta) - 75 = 90 + \theta - 75 = 15 + \theta \\
 \angle BOS &= 90 - \angle EOB = 75 - \theta \\
 \angle BOC &= 2(75 - \theta) = 150 - 2\theta
 \end{aligned}$$

$$\begin{aligned}
 \text{(ii)} \quad \angle HCO &= 360 - 90 - (90 + \theta) - \angle BOC \\
 &= 180 - \theta - (150 - 2\theta) \\
 &= 30 + \theta \\
 \phi &= 90 - \angle HCO = 60 - \theta \\
 \phi &= 60 - \theta
 \end{aligned}$$

$$\begin{aligned}
 \text{(iii)} \quad \text{If } \theta = \phi \quad & 2\theta = 60 \\
 & \theta = 30
 \end{aligned}$$

Total internal reflection at O requires

$$\begin{aligned}
 n &> \frac{\sin 90}{\sin(75 - \theta)} && \text{i.e. } \sin 90 / \sin BOS && 2 \\
 &> 1 / \sin 45 \\
 n &> 1.414 && && 1
 \end{aligned}$$

(iv)  $90^\circ$  as the faces EH and HG are perpendicular and the angles of refraction at these faces are equal,  $\theta$  and  $\phi$  equally inclined to the faces, in the same sense

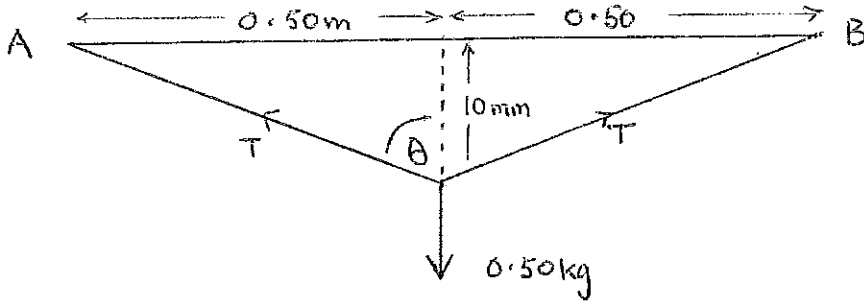
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[8]

(d)

The ripples on the water surface act like a continuous series of mirrors inclined to the horizontal with  $\pm$  a few degrees. This produces a continuous series of images of the light forming a light 'pillar' i.e. elongated image of the light source [2]

Q3



Increase in length of wire  $\Delta l$  given by, using Pythagoras' theorem,

$$\begin{aligned}\Delta l &= 2\sqrt{(5 \times 10^{-1})^2 + (10^{-2})^2} - 1.00 \\ &= 2(5.0 \times 10^{-1}) \sqrt{1 + \left(\frac{10^{-2}}{5 \times 10^{-1}}\right)^2} - 1.00 \\ &= \sqrt{1 + \left(\frac{1}{50}\right)^2} - 1.00 \\ &= \frac{1}{2} \left(\frac{1}{50}\right)^2 \dots \quad \text{using Binomial Th or directly} \\ &= 2.0 \times 10^{-4} \text{ m}\end{aligned}$$

$$\therefore \frac{\Delta l}{l} = 2.0 \times 10^{-4}$$

Resolving forces

$$2T \cos \theta = 0.50 (9.81)$$

$$2T \frac{10}{\sqrt{(500)^2 + (10)^2}} = 0.50 (9.81)$$

$$2T = (0.50)(9.81) \sqrt{(50)^2 + 1}$$

$$T \approx (0.25)(50)(9.81)$$

$$\approx 122.6 \text{ N}$$

Thus Young's modulus

$$Y = \frac{\frac{122.6}{10^{-6}}}{2.0 \times 10^{-4}}$$

$$Y = 6.1 \times 10^{11} \text{ Nm}^{-2}$$

[10]

Q3

i) (i)

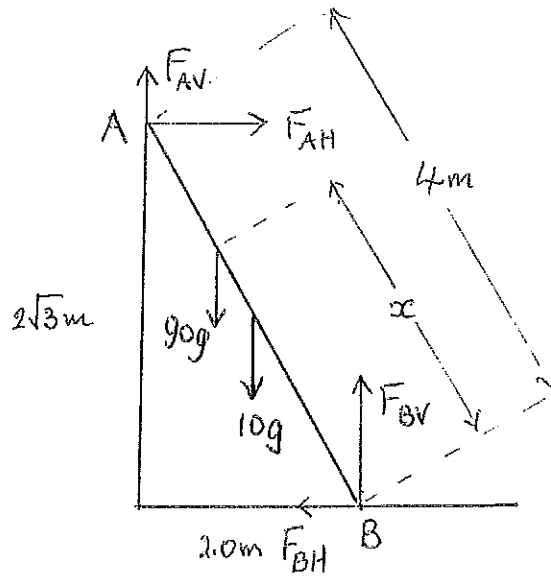


Diagram 2 MARKS ; take  $\frac{1}{2}$  off for any missing forces

2

(ii) For smooth wall  $F_{AV} = 0$

Resolving vertically  $F_{BV} = 90g + 10g = 100g$

Resolving horizontally  $F_{BH} = F_{AH}$

1

1

Taking moments about A

$$2\sqrt{3} F_{BH} + 1(10g) + (4-x)\frac{1}{2}(90g) - 2F_{BV} = 0$$

2

Substituting  $F_{BV} = 100g$

$$F_{BH} = \frac{[200 - 10 - (4-x)\frac{1}{2}90]g}{2\sqrt{3}}$$

$$F_{BH} = \frac{(20 + 90x)g}{4\sqrt{3}} \quad \text{①}$$

2

If  $\mu$  coeff. of friction...

$$\begin{aligned} \text{Now } F_{BH} &\leq \mu F_{BV} \\ &\leq 100\mu g \end{aligned}$$

Thus from ①

$$\mu \geq \frac{20 + 90x}{100(4\sqrt{3})}$$

For  $x = 4.0m$

$$\mu \geq 0.55$$

Minimum value of  $\mu$  is 0.55

2

Q4

(i) At equilibrium, with spring constant  $k$ ,  
 $kx_1 = Mg$ , where  $\omega^2 = \frac{k}{M}$  1

Giving

$$x_1 = \frac{Mg}{k} = \frac{Mg}{M\omega^2}$$

$$\underline{x_1 = \frac{g}{\omega^2}}$$

[2]

(ii)

$$\begin{aligned} V &= \frac{1}{2}k(x+x_1)^2 - Mgx \\ &= \frac{1}{2}M\omega^2(x+x_1)^2 - Mgx \\ &= \frac{1}{2}M\omega^2\left(x + \frac{g}{\omega^2}\right)^2 - Mgx \\ &= \frac{1}{2}M\omega^2\left(x^2 + \frac{2gx}{\omega^2} + \left(\frac{g}{\omega}\right)^2\right) - Mgx \\ \underline{V} &= \frac{1}{2}M\omega^2x^2 + \frac{1}{2}M\left(\frac{g}{\omega}\right)^2 \quad \text{(A)} \end{aligned}$$

} 1

[3]

$V$  depends on  $x^2$ , not  $x$ , and  $\frac{1}{2}M\left(\frac{g}{\omega}\right)^2$  is a constant

(iii) Max PE occurs when  $x=A$ , so  $V_{\max}$  given by  
 $V_{\max} = \frac{1}{2}M\omega^2A_1^2 + \frac{1}{2}M\left(\frac{g}{\omega}\right)^2$

This occurs when velocity stationary. Change in PE from  $x=A_1$  to  $x=0$  is from (A)

$$\frac{1}{2}M\omega^2A_1^2$$

This is the KE at equilibrium position:  $T_1 = \frac{1}{2}M\omega^2A_1^2$   
 (velocity  $\omega A_1$ )

Thus

$$\underline{E_1 = V_{\max} = \frac{1}{2}M\omega^2A_1^2 + \frac{1}{2}M\left(\frac{g}{\omega}\right)^2}$$

[2]

(iv) If after collision  $M$  moving with velocity  $V_1$  and  $\frac{1}{2}M$  has velocity  $V_2$   
 Conservation of momentum

$$M(A_1\omega) = MV_1 + \frac{1}{2}MV_2 \quad \text{ie } A_1\omega = V_1 + \frac{1}{2}V_2 \quad 1$$

Conservation of energy

$$\frac{1}{2}M(A_1\omega)^2 = \frac{1}{2}MV_1^2 + \frac{1}{2}MV_2^2 \quad \text{ie } (A_1\omega)^2 = V_1^2 + \frac{1}{2}V_2^2 \quad 1$$

From these equations

$$(A_1\omega)^2 = V_1^2 + \frac{1}{2}4(A_1\omega - V_1)^2$$

thus

$$3V_1^2 - 4V_1(A_1\omega) + (A_1\omega)^2 = 0$$

24

(iv) Solving, 
$$v_1 = \frac{1}{6} [4(A_1\omega) \pm \sqrt{16-12} (A_1\omega)]$$

$$v_1 = \frac{1}{6} (4 \pm 2) A_1\omega$$

$$v_1 = A_1\omega \text{ or } \frac{1}{3} A_1\omega$$
} 2

$v_1 = \frac{1}{3} A_1\omega$  only acceptable solution if changed on collision  
 Thus amplitude reduces to  $A_2 = \frac{1}{3} A_1$ , as  $vel^2 = amp^2 \times \omega^2$

[5]

(v) New equilibrium position,  $x_2$ , given by

$$kx_2 = \frac{3}{2} Mg$$

$$x_2 = \frac{3g}{2\omega^2}$$

as  $k = M\omega^2$

Conservation of momentum gives new velocity after collision,  $u$ ,

$$M(A_1\omega) = \frac{3}{2} Mu_2$$

$$u_2 = \frac{2}{3} A_1\omega$$

Potential energy using (a), with  $M$  replaced by  $(\frac{3}{2}M)$  and

$x_2 - x_1 = \frac{g}{2\omega^2}$ , measured from new equilibrium position

$$P.E. = \frac{1}{2} (\frac{3}{2}M) (\frac{g}{2\omega^2})^2 + \frac{1}{2} (\frac{3}{2}M) (\frac{g}{\omega})^2$$

The KE after collision is  $\frac{1}{2} (\frac{3}{2}M) u_2^2 = \frac{1}{2} (\frac{3}{2}M) (\frac{2}{3} A_1\omega)^2$

Equating total energy at  $t=0$  velocity, amplitude  $A_2$ , and following collision, measured from new equilibrium position

$$\frac{1}{2} (\frac{3}{2}M) \omega^2 A_2^2 + \frac{1}{2} (\frac{3}{2}M) (\frac{g}{\omega})^2 = \frac{1}{2} (\frac{3}{2}M) (\frac{2}{3} A_1\omega)^2 + \frac{1}{2} (\frac{3}{2}M) \omega^2 (\frac{g}{2\omega^2})^2 + \frac{1}{2} (\frac{3}{2}M) (\frac{g}{\omega})^2$$

$$\frac{1}{2} \omega^2 A_2^2 = \frac{1}{2} \omega^2 (\frac{g}{2\omega^2})^2 + \frac{1}{2} (\frac{2}{3} A_1\omega)^2$$

$$A_2 = \left[ \left( \frac{g}{2\omega^2} \right)^2 + \left( \frac{2A_1}{3} \right)^2 \right]^{1/2}$$

[8]



Q5

(a) (i)  $Q = CV$

For  $2.0 \mu\text{F}$  capacitor:  $120 \times 10^{-6} = 2.0 \times 10^{-6} V_1$

$V_1 = 60V$

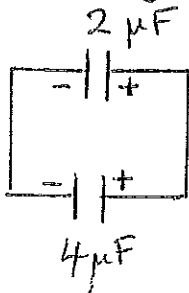
$\frac{1}{2}$

For  $4.0 \mu\text{F}$  capacitor:  $120 \times 10^{-6} = 4.0 \times 10^{-6} V_2$

$V_2 = 30V$

$\frac{1}{2}$

Connecting capacitors



If charge  $Q$  on  $2 \mu\text{F}$  capacitor and  $240 - Q$  on  $4 \mu\text{F}$  capacitor (conservation of charge) common potential  $V_c$  given by

$V_c = \frac{Q}{2.0 \times 10^{-6}} = \frac{240 - Q}{4.0 \times 10^{-6}}$

1

Giving

$Q = 80 \mu\text{C}$

and  $V_c = 40V$

1

(ii) Energy of a capacitor =  $\frac{1}{2} CV^2 = \frac{1}{2} QV$

$\frac{1}{2}$

Initial total energy =  $\frac{1}{2} (120 \times 10^{-6}) (60 + 30) = 5.4 \times 10^{-3} \text{ J}$

1

Final energy =  $\frac{1}{2} (40) (80 + 160) 10^{-6} = 4.8 \times 10^{-3} \text{ J}$

1

Energy lost due to heating of  $6 \times 10^{-4} \text{ J}$

$\frac{1}{2}$

6

Q5

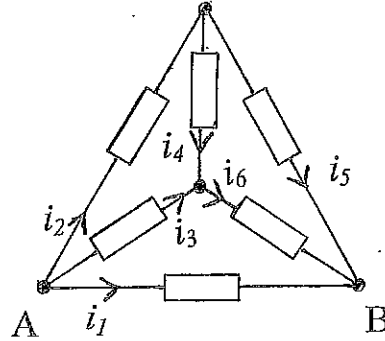
b) (i) The currents are all reversed

When  $V$  replaced by  $-V$  the circuit is just a 'reflection' of the original circuit. So comparing currents in the original circuit and its reflection

$$i_2 = i_6$$

$$i_3 = i_5$$

$$i_4 = -i_4$$



(ii)  $i_1 = \frac{V}{R}$

As  $i_4 = -i_4$ ,  $i_4 = 0$

So  $i_2 = i_6 = \frac{V}{2R}$

and  $i_3 = i_5 = \frac{V}{2R}$

(iii) As  $i_4 = 0$ ,  $2R, 2R$  and  $R$  are all in parallel with each other

$$\text{Total resistance across AB} = \left( \frac{1}{2R} + \frac{1}{2R} + \frac{1}{R} \right)^{-1}$$

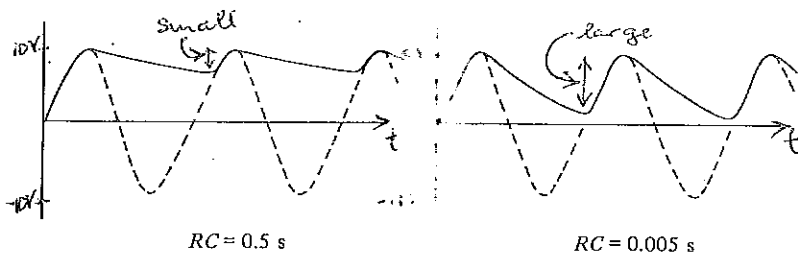
$$= \frac{R}{2}$$

(c) (i) Time constant =  $CR$

For  $10k\Omega$   $CR = (50 \times 10^{-6}) (10^4) = 0.5s$

$CR = (50 \times 10^{-6}) (100) = 5.0 \times 10^{-3}s$

(ii)



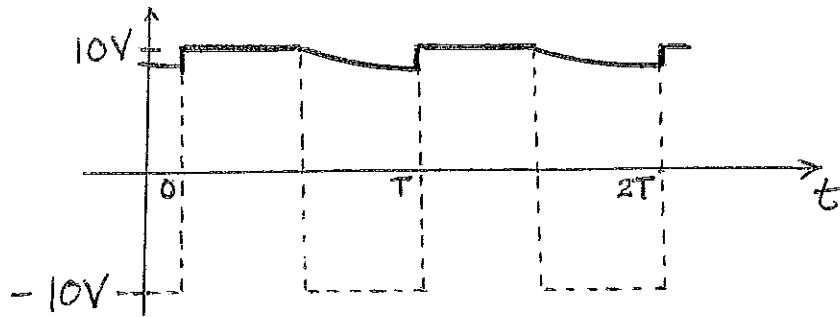
1+1 = 2  
One mark for each graph correct

1/2  
1/2  
1/2  
1/2  
1  
1

1  
1/2  
7  
1/2  
1/2

5(c)

(iii)



Capacitor will discharge for half a period,  $\frac{T}{2}$

$$\text{i.e. } \frac{1}{2} \left( \frac{1}{50} \right) = 10^{-2} \text{ s}$$

Voltage at time  $t$ ,  $V(t) = 10 e^{-t/RC}$

At  $t = 10^{-2} \text{ s}$ ,

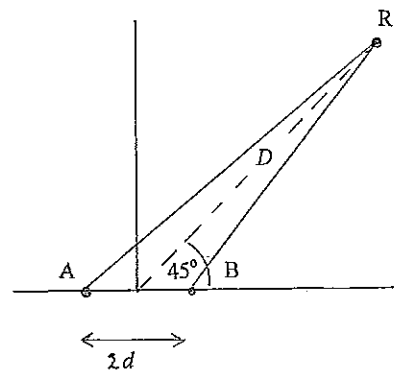
$$\begin{aligned} V &= 10 \exp\left(-\frac{10^{-2}}{0.5}\right) \\ &= 10 \exp(-0.02) \\ &= 9.782 \text{ V} \end{aligned}$$

$$\text{Fractional change} = \frac{0.218}{10} \approx \underline{\underline{0.022}}$$

[7]

Q6

Using cosine rule in diagram,



$$(AR)^2 = D^2 + \left(\frac{1}{4}\right)^2 - 2D\left(\frac{1}{4}\right)\cos 135^\circ$$

$$= D^2 + \left(\frac{1}{4}\right)^2 + \frac{1}{2}D\left(\frac{1}{2}\right)$$

$$(AR)^2 = D^2 + \left(\frac{1}{4}\right)^2 + \frac{1}{4}\sqrt{2}D$$

$$(BR)^2 = D^2 + \left(\frac{1}{4}\right)^2 - 2\left(\frac{1}{4}\right)D\cos 45^\circ$$

$$(BR)^2 = D^2 + \left(\frac{1}{4}\right)^2 - \frac{1}{4}\sqrt{2}D$$

$$(i) \quad t_A = \frac{AR}{c_s} = \frac{1}{c_s} \left[ D^2 + \left(\frac{1}{4}\right)^2 + \frac{1}{4}\sqrt{2}D \right]^{\frac{1}{2}} \quad (1)$$

$$(ii) \quad t_B = \frac{BR}{c_s} = \frac{1}{c_s} \left[ D^2 + \left(\frac{1}{4}\right)^2 - \frac{1}{4}\sqrt{2}D \right]^{\frac{1}{2}} \quad (2)$$

$$\frac{1}{[4]}$$

b) Frequency emitted at time  $t$  is  $f = 10t$

(i) Frequency reaching R from A will have been emitted at  $(t - t_A)$   
 Thus " " " " "  $f_A = 10(t - t_A)$  ( $t_A$  given by (1))

(ii) Similarly for B,  $f_B = 10(t - t_B)$  ( $t_B$  given by (2))

$$\frac{1}{[4]}$$

c) As  $D \gg \frac{1}{4} \text{ m}$  in (1) and (2),  $AR$  and  $BR$  are approximately equal to  $D$ . Thus  $f_A$  and  $f_B$  are approximately equal to, say,  $f_{AB}$  where

$$f_{AB} = 10\left(t - \frac{D}{c_s}\right)$$

The error in 'D' term

$$\begin{aligned} \Delta D &= \pm d \cos 45^\circ \\ &= \pm \frac{1}{4} \frac{\sqrt{2}}{2} \text{ m} \\ &= \pm \frac{1}{8\sqrt{2}} \text{ m} \end{aligned}$$

Thus accuracy in  $f_{AB}$

$$\begin{aligned} \Delta f &= \pm 10 \left(\frac{\sqrt{2}}{8}\right) \frac{1}{c_s} \\ \Delta f &= \pm 1.077 \left(\frac{1}{c_s}\right) \text{ Hz} \end{aligned}$$

$$\frac{1}{4}$$

(d) Wavelength  $\lambda$  given by

$$\lambda = \frac{c_s}{f_{AB}} = \frac{c_s}{10\left(t - \frac{D}{c_s}\right)} \quad (3)$$

1

Q6

d) For the first minimum the path difference of the sound reaching R from A and B must differ by  $\lambda/2$ .  
So

$$2d \cos 45^\circ = \frac{1}{2} \lambda$$

or

$$4d \left( \frac{\sqrt{2}}{2} \right) = \lambda = \frac{c_s}{10(t - D/c_s)} \quad \text{from (3)}$$
$$4 \left( \frac{1}{4} \right) \left( \frac{\sqrt{2}}{2} \right) 10 \left( t - \frac{D}{c_s} \right) = c_s$$

As  $t = 52.2 \text{ s}$  and  $D = 1200\sqrt{2} \text{ m}$ ,

$$10 \left( 52.2 - \frac{1200\sqrt{2}}{c_s} \right) \frac{1}{\sqrt{2}} = c_s$$
$$396.01 c_s - 1200 = c_s^2$$
$$c_s^2 - 396.01 c_s + 1200 = 0$$

Solving

$$c_s = \frac{1}{2} \left[ 396.01 \pm \sqrt{(396.01)^2 - 4(1200)} \right]$$
$$= 363 \text{ ms}^{-1} \quad \text{or} \quad 33 \text{ ms}^{-1}$$

Only physically acceptable solution is  $c_s = 363 \text{ ms}^{-1}$

[7]

Q7

(a) If  $N_0$  be the number of atoms initially present and  $t$  years the age of the Earth  
 For  $U^{238}$

$$N_1(t) = N_0 e^{-\alpha t}$$

For  $U^{235}$

$$N_2(t) = N_0 e^{-\beta t}$$

Then

$$\frac{N_1}{N_2} = 140 = e^{-(\alpha - \beta)t}$$

$$\ln(140) = (\beta - \alpha)t = \ln 2 \left( \frac{1}{7.13 \times 10^8} - \frac{1}{4.05 \times 10^9} \right) t$$

$$\frac{\ln(140)}{\ln(2)} = \frac{1}{10^9} \left( \frac{10}{7.13} - \frac{1}{4.05} \right) t$$

$$7.129 = 10^{-9} (1.4025 - 0.2222) t$$

$$7.129 = 10^{-9} (1.180) t$$

$$t = 6.04 \times 10^9 \text{ years}$$

b)

$$N = N_0 e^{-\lambda t}$$

$$N(30) = 4 \times 10^5 e^{-30\lambda}$$

$$= (4 \times 10^5) \exp\left(-\frac{30}{45} \ln 2\right)$$

$$= (4 \times 10^5) \exp\left(-\frac{2}{3} \ln 2\right)$$

$$= (4 \times 10^5) \exp\left(-\ln 2^{2/3}\right)$$

$$= (4 \times 10^5) 2^{-2/3}$$

$$N(30) = 2.52 \times 10^5 \text{ Bq}$$

here  $\lambda = \ln 2 / 45 \text{ days}^{-1}$

FRACTION OF MASS REMOVED

Activity in  $5 \times 10^{-3} \text{ m}^3$  of oil

$$\frac{126}{10 \times 60} \left( \frac{5.0 \times 10^{-3}}{10^3 \times 10^{-6}} \right) = \frac{126}{600} (50) = \frac{21}{2} \text{ Bq}$$

$$\text{Fraction of mass removed} = \frac{\frac{21}{2}}{2.52 \times 10^5} = \frac{4.02 \times 10^{-5}}{2.52 \times 10^5}$$

$$\text{ie } 4.02 \times 10^{-3} \%$$

6

c)  $N = N_0 e^{-\lambda t}$  (A)  $\frac{1}{2}$

where

$$\lambda = \ln 2 / 70 = 0.00990 \text{ s}^{-1}$$
 $\frac{1}{2}$

Substituting  $N_0 = 82$  into RHS of A gives

$$82 \exp(-210(0.00990)) = 10.3 \neq 19$$

Thus  $N = 19$  and  $N_0 = 82$  does not satisfy (A)

If there is a constant background of  $k$  counts  $\text{s}^{-1}$  then (A) gives

$$(19 - k) = (82 - k) \exp[-210(0.00990)]$$

Giving  $\frac{19 - k}{82 - k} = 0.125$

$$\underline{k = 10 \text{ counts s}^{-1}}$$
[5]

(d) Here  $N = N_1 e^{-\lambda_1 t} + N_2 e^{-\lambda_2 t}$   $\lambda_2 \ll \lambda_1$

For times  $t \gg T_1$  the main source of radioactivity would be due to the second source. Then

$$N \approx N_2 e^{-\lambda_2 t}$$
 $\text{half-lives } T_2 \gg T_1$

Taking measurements and plotting  $\ln N$  against  $t$  will give  $\lambda_2$  and hence  $T_2 = \ln 2 / \lambda_2$ .

[3]

Q8

a)

$$g = \frac{GM_E}{R_E^2} = \frac{(6.67 \times 10^{-11})(5.97 \times 10^{24})}{(6.38 \times 10^6)^2} = G(1.4667 \times 10^{11})$$

$$g = 9.78 \text{ ms}^{-2}$$

1  
1  

---

[3]

b)

1ST SOLUTION

$$g = \frac{GM}{R^2}$$

$$\Delta g = \frac{G\Delta M}{R^2} + (-2) \frac{GM\Delta R}{R^3}$$

$$\frac{\Delta g}{g} = \frac{\Delta M}{M} - \frac{2\Delta R}{R}$$

Now

$$\frac{\Delta R}{R} = \frac{-15 \times 10^3}{6.38 \times 10^6}$$

and

$$\frac{\Delta M}{M} = \frac{-4\pi (6.38 \times 10^6)^2 (15 \times 10^3) (2.7 \times 10^3)}{5.97 \times 10^{24}}$$

$$= -2.35 \times 10^{-3}$$

$$= -3.47 \times 10^{-3}$$

Thus

$$\frac{\Delta g}{g} = -3.47 \times 10^{-3} + 4.70 \times 10^{-3}$$

$$\frac{\Delta g}{g} = +1.23 \times 10^{-3}$$

(one of the marks for noting increase in g, +ve sign, one mark for numerical value)

1  
2  

---

7

2ND ALTERNATIVE SOLUTION

$$g = \frac{GM_E}{R^2} = 1.4667 \times 10^{11} G \quad \text{①}$$

At 15 km  $g_{15} = \frac{G[5.97 \times 10^{24} - 4\pi (6.38 \times 10^6)^2 (15 \times 10^3) (2.7 \times 10^3)]}{(6.38 \times 10^6 - 15 \times 10^3)^2}$

$$= \frac{G[5.97 \times 10^{24} - 2.0716 \times 10^{22}]}{(6.365 \times 10^6)^2}$$

$$= \frac{G \cdot 5.949 \times 10^{24}}{40.51 \times 10^{12}}$$

$$= G(1.4684 \times 10^{11})$$

From ①

$$\Delta g = + G(1.8 \times 10^8)$$

$$\frac{\Delta g}{g} = \frac{1.8 \times 10^8}{1.47 \times 10^{11}} = +1.23 \times 10^{-3}$$

(mark for +ve sign, increase in g, 1 mark for magnitude)

2  

---

[7]



Q8

c)

$$\Delta g = \frac{G \frac{4}{3}\pi (2.5 \times 10^3)^3 (7.9 - 2.7) 10^3}{(2.5 \times 10^3)^2}$$

3

$$= G \frac{4}{3}\pi (2.5 \times 10^3) (5.2) 10^3$$

$$= G (5.4 \times 10^7)$$

$$= 6.67 \times 10^{-11} (5.4 \times 10^7)$$

$$= 3.63 \times 10^{-3}$$

$$\frac{\Delta g}{g} = +3.71 \times 10^{-4}$$

magnitude 1 mark  
sign +, increasing  $\frac{1}{2}$  mark

1 1/2

1 1/2

[6]

d)

If spherical volume of iron removed

$$\Delta g = - \frac{G \left(\frac{4}{3}\pi\right) (2.5 \times 10^3)^3 7.9}{(2.5 \times 10^3)^2}$$

$$= - G \left(\frac{4}{3}\pi\right) (2.5 \times 10^3) 7.9$$

$$= - (6.67 \times 10^{-11}) \left(\frac{4}{3}\pi\right) (2.5 \times 10^3) 7.9$$

$$= - 5.52 \times 10^{-6}$$

$$\frac{\Delta g}{g} = - 5.64 \times 10^{-7}$$

$\frac{1}{2}$  mark for -ve sign  
1 mark magnitude

1 1/2

1

1 1/2

[4]

(a)

(i)

$$\frac{GM_E}{(R_{EM} - R_{MF})^2} = \frac{GM_M}{R_{MF}^2}$$

$$\frac{R_{EM} - R_{MF}}{R_{MF}} = \left(\frac{M_E}{M_M}\right)^{\frac{1}{2}}$$

$$= \left(\frac{5.97 \times 10^{24}}{7.35 \times 10^{22}}\right)^{\frac{1}{2}}$$

$$= 9.01$$

$$R_{EM} - R_{MF} = 9.01 R_{MF}$$

$$10.01 R_{MF} = R_{EM} = 3.84 \times 10^5 \text{ km}$$

$$\underline{R_{MF} = 3.84 \times 10^4 \text{ km}}$$

(ii)

$$\frac{GM_E}{R_{EM} - R_{MP}} = \frac{GM_M}{R_{MP}}$$

$$\frac{R_{EM} - R_{MP}}{R_{MP}} = \frac{M_E}{M_M}$$

$$= \frac{5.97 \times 10^{24}}{7.35 \times 10^{22}} = 8.12 \times 10$$

$$R_{EM} = R_{MP} (82.2) = 3.84 \times 10^5$$

$$\underline{R_{MP} = 4.67 \times 10^3 \text{ km}}$$

[6]

(b)

Let  $v$  be the velocity of impact on lunar surface and  $M_R$  the mass of the rocket. At  $R_{MF}$  rocket has zero velocity if it has the minimum energy to enable it to reach the Moon. Conservation of energy requires

$$-GM_R M_M \left[\frac{1}{R_{MF}}\right] - GM_R M_E \left[\frac{1}{R_{EM} - R_{MF}}\right] = \frac{1}{2} M_R v^2 - GM_R M_M \left[\frac{1}{R_M}\right] - GM_R M_E \left[\frac{1}{R_{EM} - R_M}\right]$$

$$\underline{v = \sqrt{2G \left[ M_M \left( \frac{1}{R_M} - \frac{1}{R_{MF}} \right) + M_E \left( \frac{1}{R_{EM} - R_M} - \frac{1}{R_{EM} - R_{MF}} \right) \right]}}$$

5

(c) For a stable orbit radius  $r$ , velocity  $v_s$ , of satellite mass  $m_s$  (neglecting influence of Earth's gravitational field)

$$\frac{m_s v_s^2}{r} = \frac{G m_s M_M}{r^2} \quad 2$$

$$v_s = \left( \frac{G M_M}{r} \right)^{\frac{1}{2}}$$

$$= \left[ \frac{(6.67 \times 10^{-11})(7.35 \times 10^{22})}{(10 + 1.74 \times 10^3) 10^3} \right]^{\frac{1}{2}}$$

$$= \left( \frac{4.90 \times 10^{12}}{1.75 \times 10^6} \right)^{\frac{1}{2}}$$

$$v_s = 1.67 \text{ km s}^{-1}$$

} 1  
1  
[4]

(d) Neglecting the influence of the Earth's gravitational field  
(d) Orbital KE,  $T_0$ , is given by

$$T_0 = \frac{1}{2} m_s v_s^2 = \frac{1}{2} (1000) (1.67 \times 10^3)^2$$

$$T_0 = 1.40 \times 10^9 \text{ J}$$

1

PE<sub>g</sub> lost in descending to Moon's surface, that increases the KE,

$$V_D = G M_M m_s \left[ \frac{1}{1.74 \times 10^6} - \frac{1}{1.75 \times 10^6} \right]$$

$$= (6.67 \times 10^{-11})(7.35 \times 10^{22})(10^3) \left( \frac{10^4}{(1.74)(1.75) 10^{12}} \right)$$

$$= 4.90 \times 10^{15} (3.28 \times 10^{-9})$$

$$= 1.61 \times 10^7 \text{ J}$$

1

1

1

Thus to 2 sig. figs energy to be extracted

is

$$E_{\text{min}} = 1.42 \times 10^9 \text{ J}$$

$$(= 1.40 \times 10^9 + 1.61 \times 10^7 \text{ J})$$

1

[5]